

PROSPECT – THERMAL DESIGN CHALLENGES FOR LUNAR VOLATILE EXTRACTION

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prospect

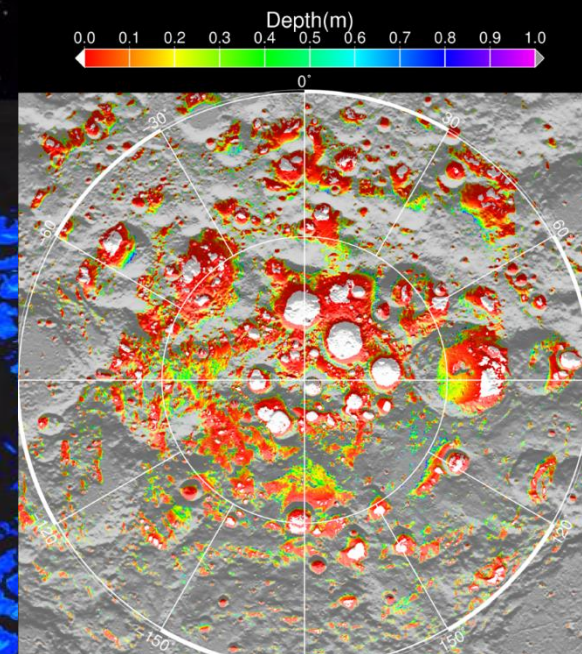


What is PROSPECT?



PROSPECT - Platform for Resource Observation and in-Situ Prospecting for Exploration, Commercial Exploitation and Transportation

- Provide first sub-surface access
 - A depths of at least 1 m
- Acquire samples from potentially volatile rich regions
 - Low-temperatures ($\sim 120\text{K}$), water-ice content, preservation
- Analyse samples for resource potential
 - thermo-chemical extraction and analysis of water, oxygen and other volatiles (identify and quantify)
- Address fundamental scientific questions
 - E.g. Isotopics and organics
- Provide samples to Russian robotic arm, for analysis in Russian instruments
 - Interface with Russian sampling tool



Calculated ice-stability* depth, based on Diviner data
*lost at a rate of $<1\text{kgm}^{-2}$ per billion years



Elements and Interfaces



Luna-27

Objective & Requirements

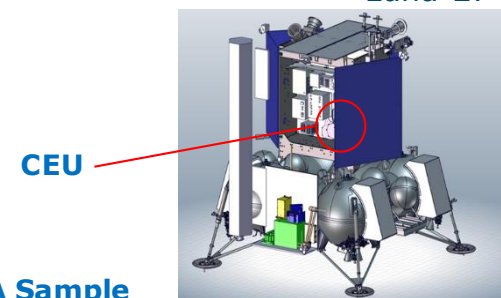
- **Find and characterize cold-trapped volatiles in lunar sub-surface samples**
- **Acquire, process & analyze** surface and **subsurface** samples during lunar day
- Landing site: 73°S to 85°S
- Operate up to 1 terrestrial year (**Survive lunar night**)
- Operational period is ~7-10 Earth days per terrestrial month
- **Sample temperature < 150 K at moment of oven sealing**
- Oven temperature **up to 1000 ° C**
- Sample transfer and gas exchange with Russian components
- Total PROSPECT **mass < 35 kg**

Elements:

- **ProSEED** Drilling and Sampling Element
- **ProSPA** Sample Analysis (including Sample Inlet System - SIS)
- Control Electronics Unit (**CEU**)
- 2 PROSPECT imagers

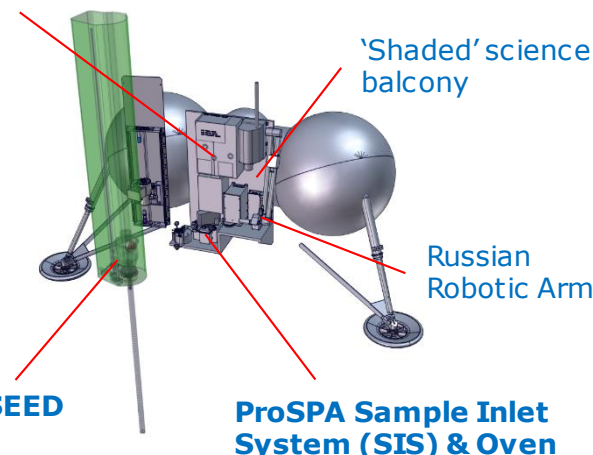
Interfaces:

- Thermo-mechanical interface with Luna-27 platform – *NPO-Lavochkin I/F*
- Electrical & Data interface via Russian BUNI (payload interface unit) – *IKI I/F*



SRR level reduced CAD model

ProSPA Sample Analysis Unit





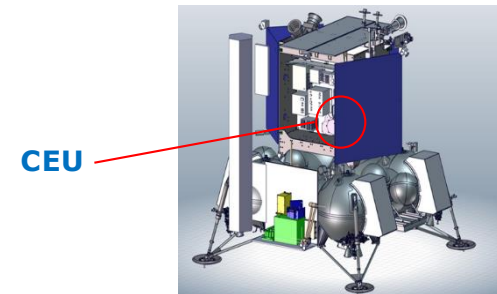
PROSPECT Thermal Requirements



Luna-27

Thermal Requirements:

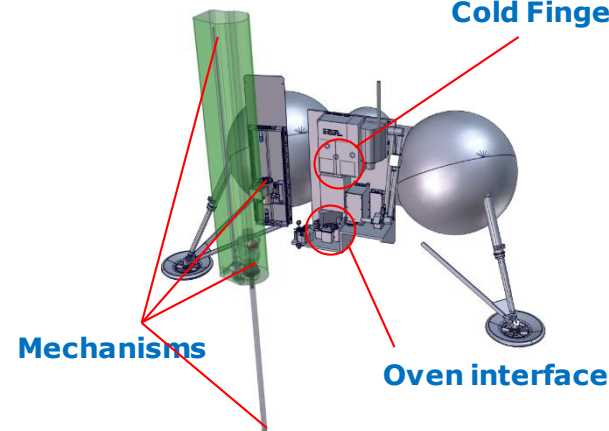
- (Some) temperature objectives:
 - Cold fingers: approx. **-100 °C**
 - Oven interface: approx. **-130 °C to -100 °C**
 - Mechanisms: approx. **-55 °C to +85 °C**
 - Increase in sample temperature **< 10 K** (Russian sample)
 - Coring, retrieval, delivery: sample temperature **< 150-170 K** (TBD)
- Interface temperatures (to lander)
 - ProSEED drill: **-180°C to +50°C**
 - PorSPA oven & SIS: **-50°C to +50°C**
 - CEU: **-50°C to +50°C**
- Lunar surface hot case:
 - IR flux **707 W/m²**, Solar flux **1425 W/m²**, 73°S plus 15° slope, Albedo moon = 0.07, Emissivity moon = 0.97
- Lunar surface cold case:
 - IR flux **0 W/m²**, Solar flux **0 W/m²**, 86°S, Albedo moon = 0.18, Emissivity moon = 0.97



CEU

SRR level reduced CAD model

Cold Fingers



Mechanisms

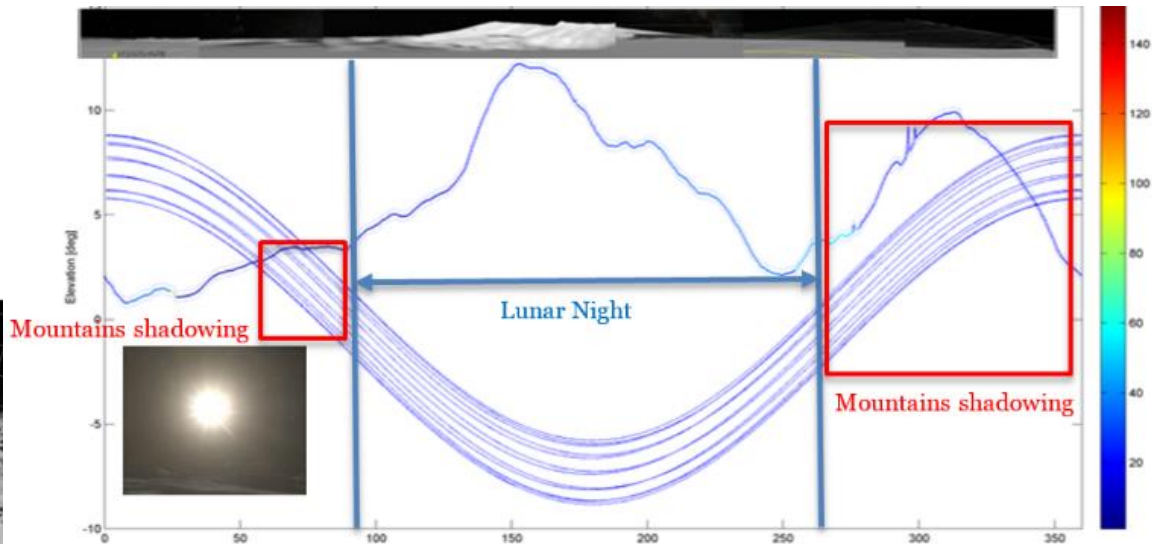
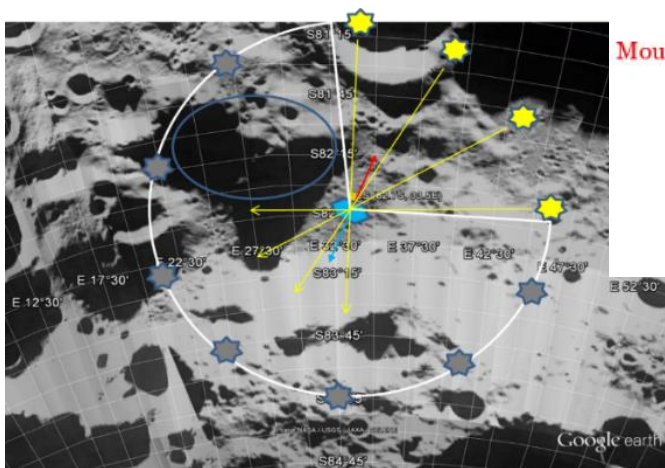
Oven interface



Investigation of environmental heat flux modelling



- Characterize environmental heat fluxes
 - Local topography
 - Sunrise & sunset
 - Surface temperature
- landing site at 82.7° S, 33.5° E



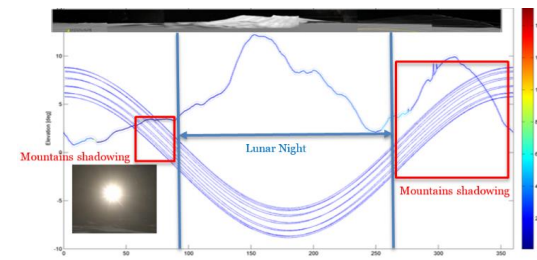
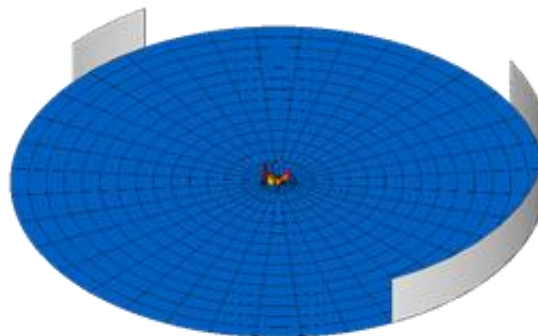


Models of the landing site



Disc Model:

- 30m radius disc with simplified horizon
- Simplified lander model
- PROSPECT components assumed to be black bodies
- Modeled in ESATAN-TMS
 - Solar Absorptivity lander = 0.6,
 - IR Emissivity lander = 0.4,

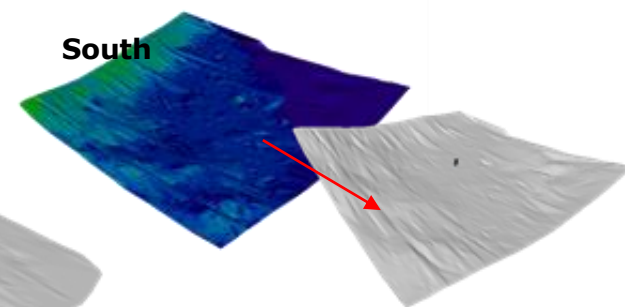
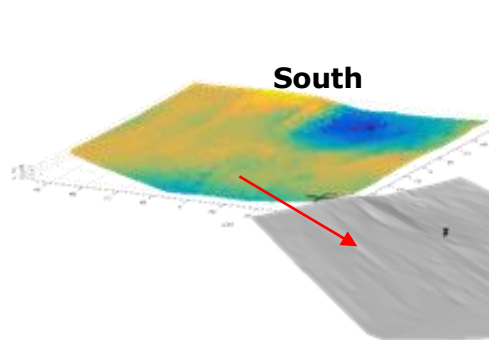


Topography Models:

- LRO LOLA data set
- 30 km x 30 km
- 1 km x 1 km
- black body cube ($T = -20^{\circ}\text{C}$)

1 km x 1 km topography model

30 km x 30 km topography model



Common assumptions:

- Thermal capacity of soil neglected
- Thermal conductivity of soil neglected
- Optical surf. prop. regolith based on literature
 - Solar Absorptivity moon = 0.93
 - IR Emissivity moon = 0.97

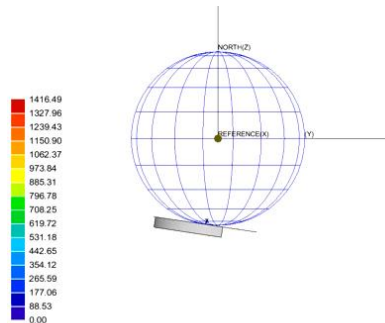




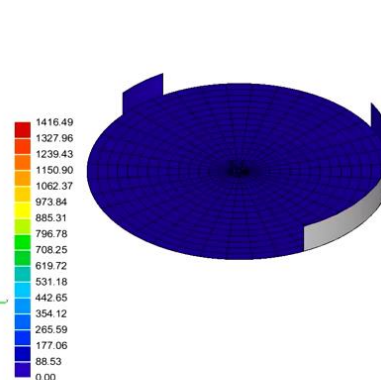
30 m Disc Model – Results



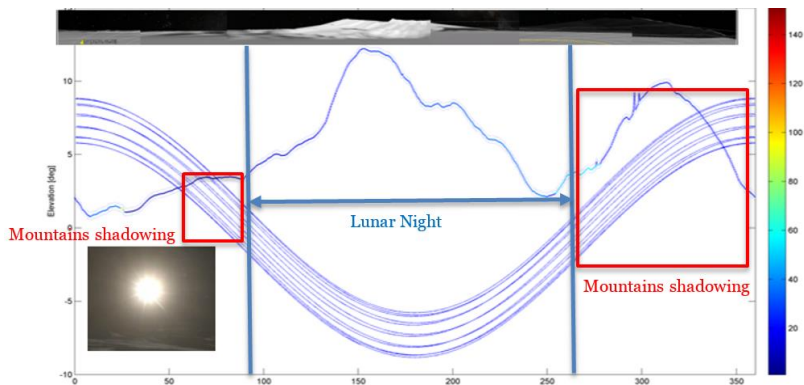
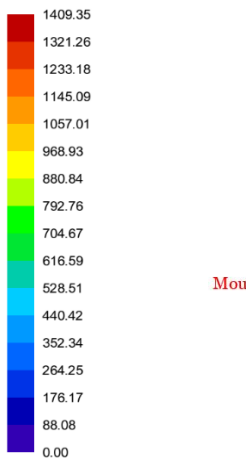
Heat flux [W/m^2]



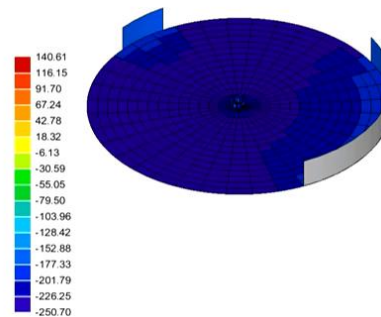
Heat flux [W/m^2]



[W/m^2]



Temperature [$^{\circ}\text{C}$]

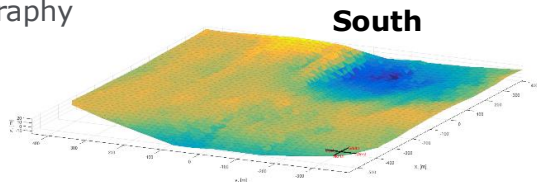




Topography Model 1 km x 1 km - Results

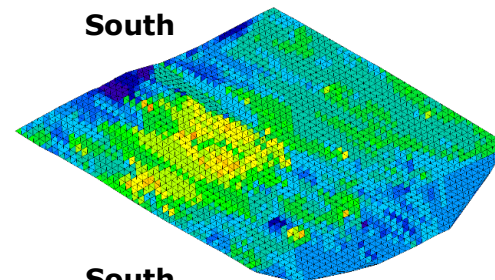


Topography



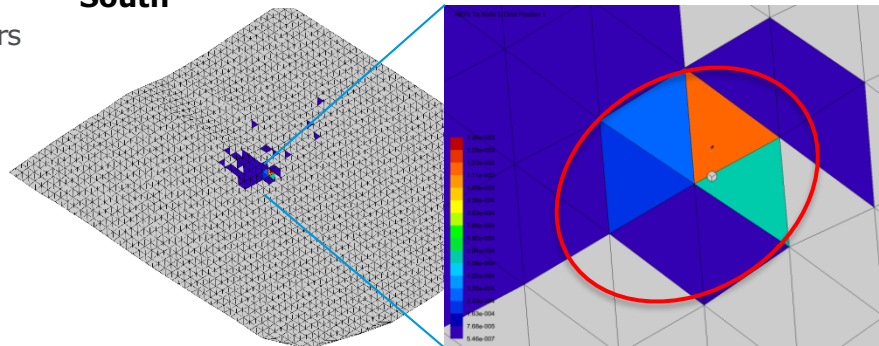
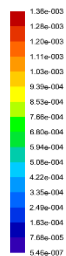
[W/m²]

South



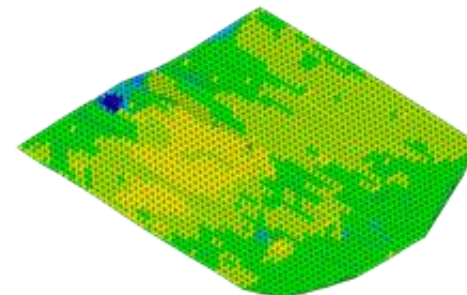
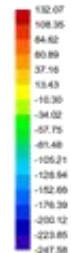
View factors

South



[°C]

South

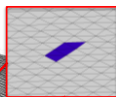
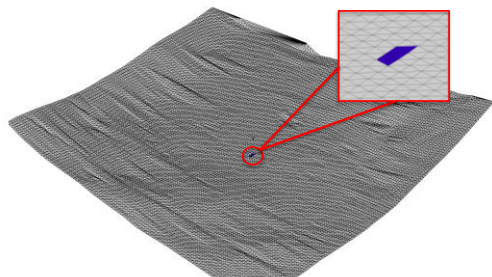


- Optical surface properties, slope angle and local shadowing effects govern surface temperature and hence environment heat fluxes
- Influencing surface elements within a 30m radius – comparable to disc model

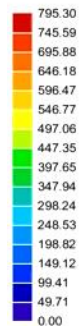




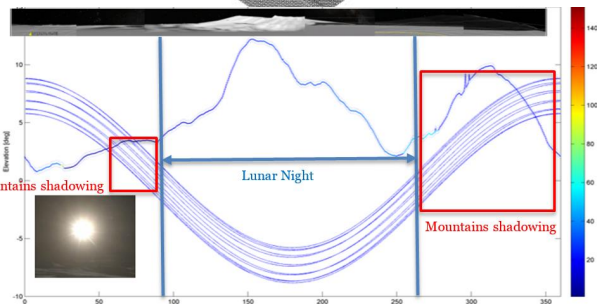
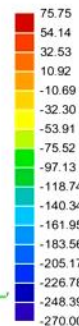
Topography Model 30km x 30km - Results



[W/m²]



[°C]





Summary & Conclusions



Summary table:

Heat Load		Surface - Disc	Surface – Topography (1 km x 1 km)	Surface – Topography (30 km x 30 km)
Solar heat load	Q_S	6.38 W	10.47 W	4.94 W
Infrared heat load	Q_{IR}	45.66 W	60.28 W	48.75 W

Note1: Heat load for 1m² area on the South facing side; no direct Sun illumination; location of PROSPECT on Luna 27

Note2: Results are specific for a lunar south pole landing site with coordinates 82.7° S, 33.5° E. Other locations will yield different results.

Conclusions:

- Feasibility of modelling the lunar surface environment with standard thermal analysis software for space applications was demonstrated
- Local / near field topography has an impact on the overall IR and solar albedo heat fluxes
- The effect of mountains located 20-30 km away from the test cube is negligible.
- 30 m disc & simplified mountains model is now used by PROSPECT industry team

Outlook (Optional):

- Re-iterate with higher resolution topographic map
- Investigate different possible landing sites
- Investigate the impact of local boulders based on to be expected size and distribution



THANK YOU

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